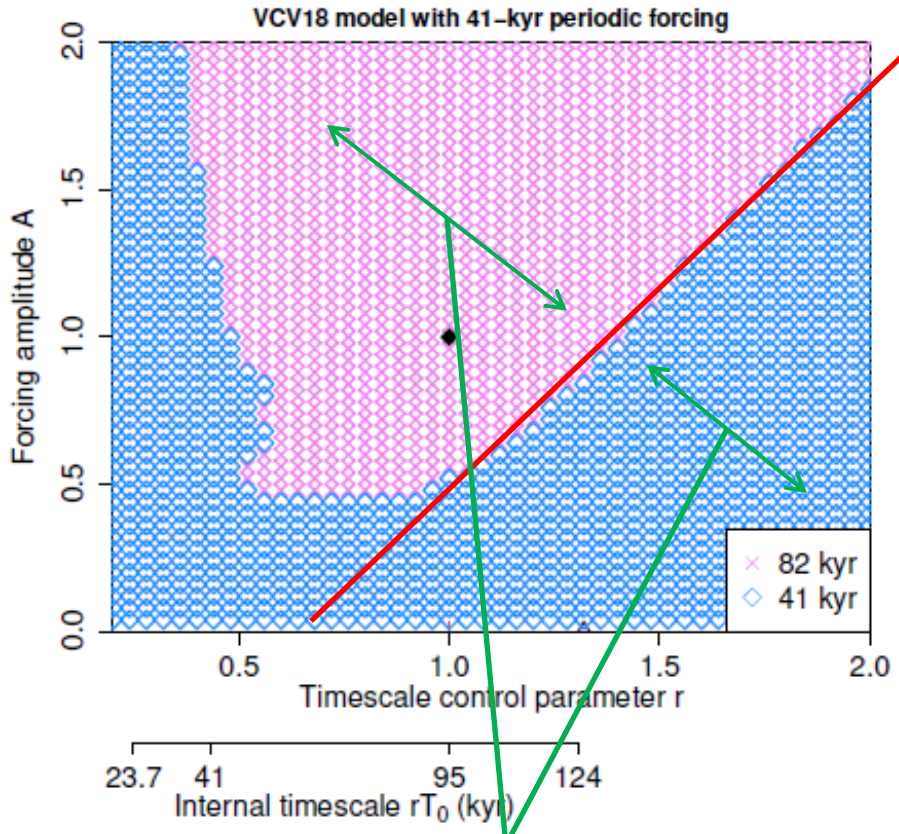

Bifurcation slope or orbitally modified intrinsic timescale $\tau_{int}\frac{a}{\varepsilon} = \frac{\zeta S_0^{1/4}}{\varepsilon} = \frac{rT_0}{A}$.

The system may bifurcate even if the intrinsic time scale is, for example, much longer than the 100-kyr eccentricity period (e.g., r =2) as long as orbital forcing is strong enough (e.g., A=2).

The system is independent on intrinsic timescale and on orbitally modified intrinsic timescale (slope).