



Supplement of

Multi-millennial-scale solar activity and its influences on continental tropical climate: empirical evidence of recurrent cosmic and terrestrial patterns

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1 **Supplementary Information (SI)**

2 **Multi-millennial-scale solar activity and its influences on continental tropical**
3 **climate: Empirical evidence of recurrent cosmic and terrestrial patterns**

4 **By J. Sánchez-Sesma**

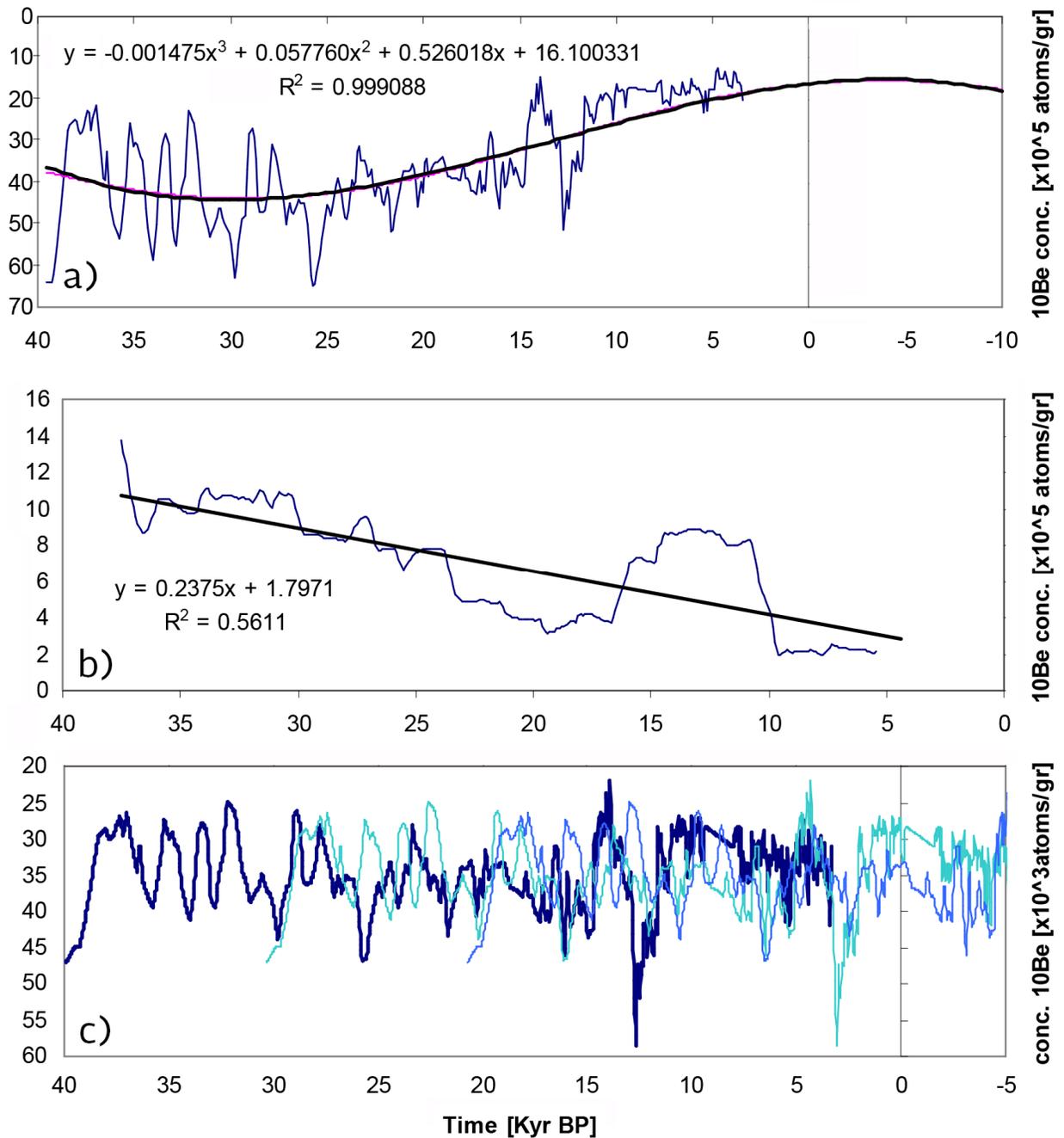
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6 **SI-1. Statistical detrending and demodulation of the solar proxy (^{10}Be) record of FN97**

7 ^{10}Be is a proxy of SA. It is produced by the impacts of galactic cosmic rays on Earth (SS14)
8 in a known process of cosmogenic nucleosynthesis. Cosmic rays are highly energetic charged
9 particles that impact Earth's upper atmosphere and terrain surface, that produce ^{10}Be , and that
10 are modulated by the variation in the strength of the geomagnetic field, as well as by solar
11 magnetic shielding (SS14).

12 In order to homogenize ^{10}Be values, we apply, firstly, a detrending process based on
13 polynomial expressions, and secondly, a demodulation intended to make the variance
14 uniform. Figure S1a shows the polynomial model for detrending, and Figure S1b shows the
15 linear model of the standard deviation for a demodulation. The results are displayed in Figure
16 S4c with two analogue models with lags of 9.6 and 19.2 Kyr, showing similar variability for
17 the next millennia.

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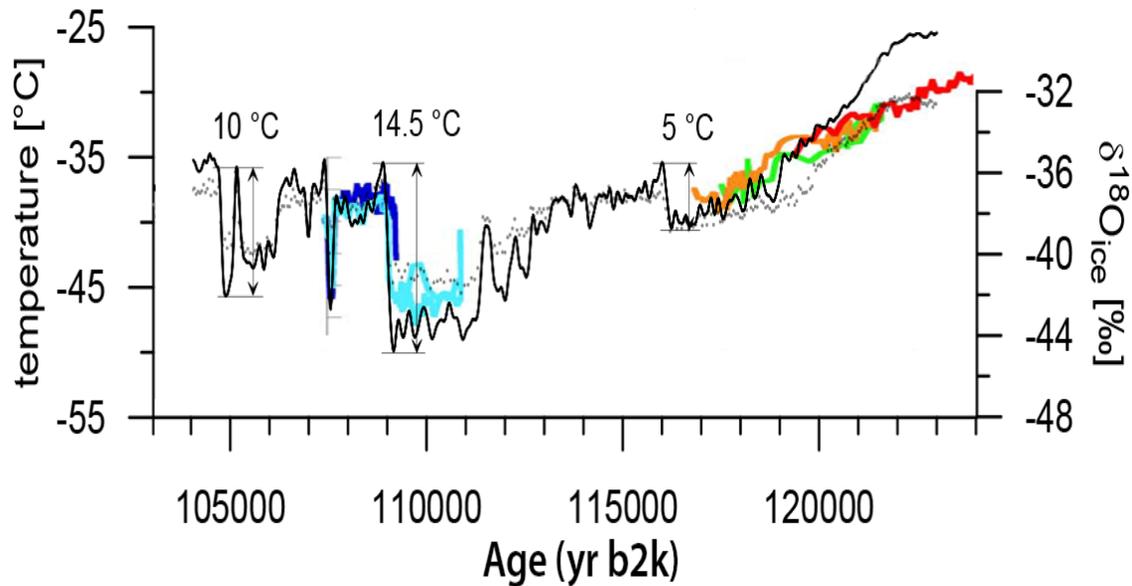


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2 Figure S1. GISP2 icecore analysis. a) 10Be concentration [N97] and a polynomial trend, b). b) Standard
 3 deviation and its linear model a). c) Demodulated 10Be [(a)/(linear model of b)], 10Be (DDR), lagged model for
 4 9.6 and 19.2 Kyr of this DDR of 10Be are also displayed. Please note that in the following figures: as the 10Be
 5 concentration varies inversely with solar activity, TSI, the beryllium scale is inverted, and thus upper parts in this
 6 scale indicate high TSI levels.

1 **SI-2. Estimation of a temporal adjustment (bias) of the Greenland 10Be record by SS14.**

2 A temporal adjustment for the NEEM 10Be record produced by SS14 is presented. This
3 record is compared with another study with information coming from the NGRIP ice-cores
4 (Kindler et al., 2014) which is located to less than 1000 km from the NEEM site. A temporal
5 bias correction of 2500 yrs going back in time was applied to SS14 data to get the best match
6 with the Kindler et al. (2014) data, and is shown in Figure S6.



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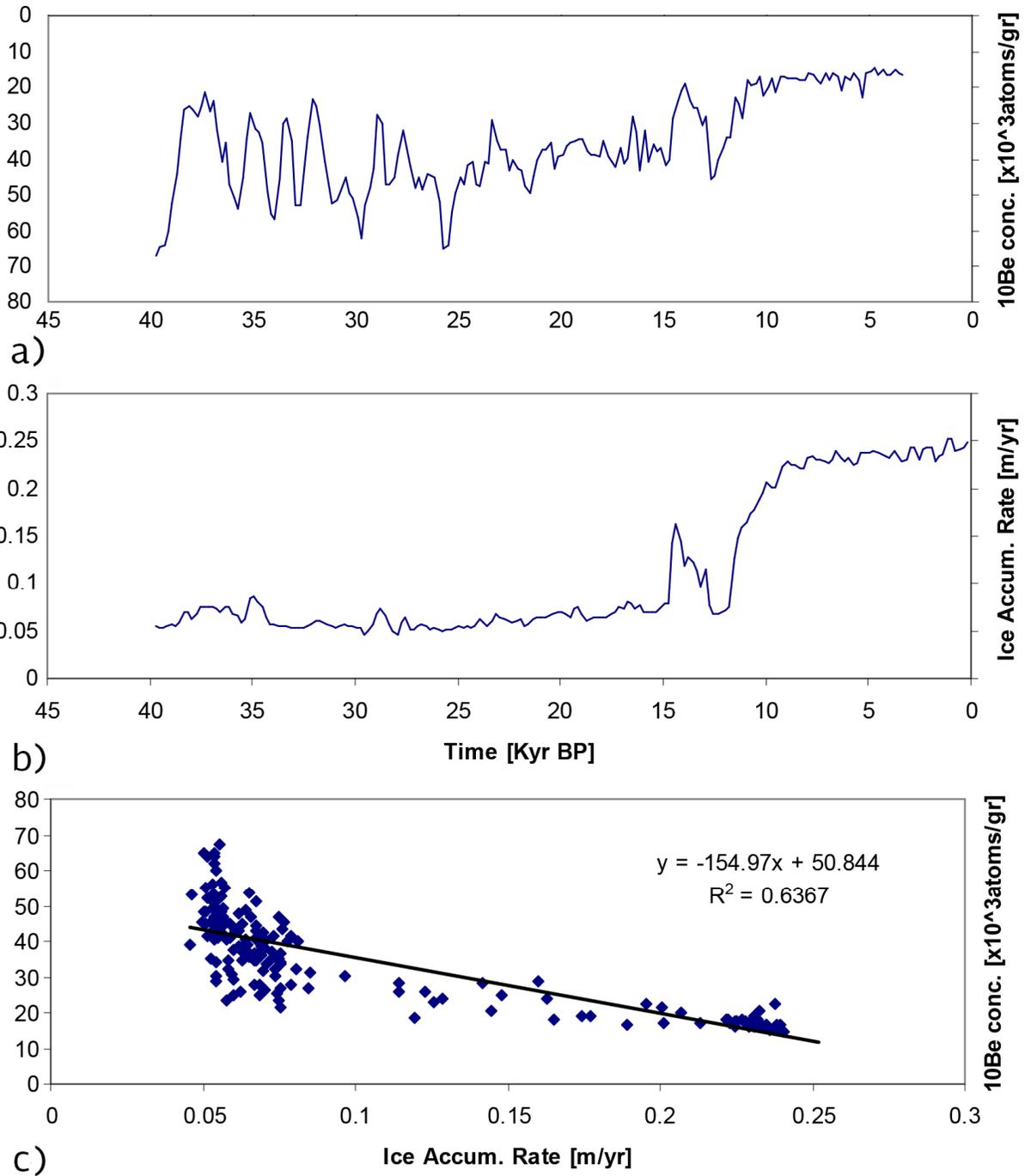
8 Figure S6. Comparison of NGRIP Greenland ice-core reconstructed isotopic anomalies in two
9 different studies (SS14, K14). The gray points indicate anomalies ¹⁸O ice and black line
10 indicates temperature from Kindler et al. (2014). Color lines are anomalies of ¹⁸O ice from
11 the SS14 that have been adjusted in time, with a lead of 2500 years to increase the match with
12 K14 results.

1 **SI-3. Detrending and demodulating the solar proxy (10Be) record of FN97, following**
2 **Alley et al. (1995)**

3 Alley et al. (1995) have proposed a method for interpolating between the end-members cases
4 of 10Be wet and dry deposition. They argued that, on average, under conditions of global
5 constant global 10Be production rate and constant dry deposition rate and scavenging ration
6 for precipitation, a linear relation should be expected between 10Be flux and snow flux at
7 GISP2. Then, deviation of measured 10Be from the correlation line can be attributed to
8 changes in the atmospheric concentration of 10Be.

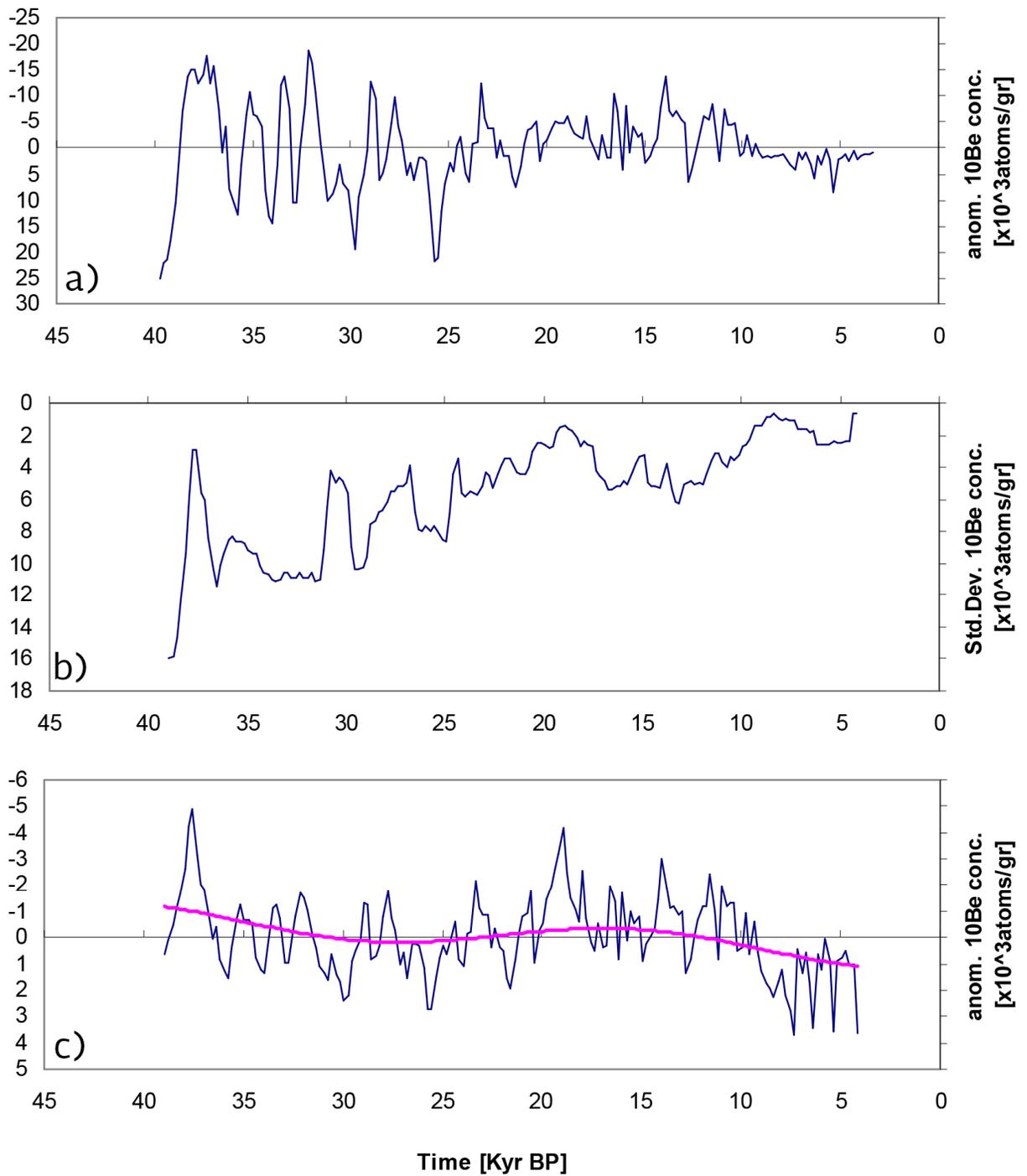
9 Figures S2 and S3 show these processes to obtain the Atmospheric 10Be. Figure S2 shows:
10 the 10Be concentration [N97], ice accumulation rate [Alley, 2000], and their linear
11 correlation. And Figure S3 shows: deviations from the linear interpolation (see S2c), its
12 standard deviation, and the demodulated deviations (S3a) with a precessional sinusoidal
13 model.

14 The final atmospheric signal of 10Be without orbital influences is displayed in Figure S4 with
15 a simple analogue model with a lag of 9.4 Kyr, showing a potential decrease of TSI for the
16 next centuries and lower values for the next millennia.



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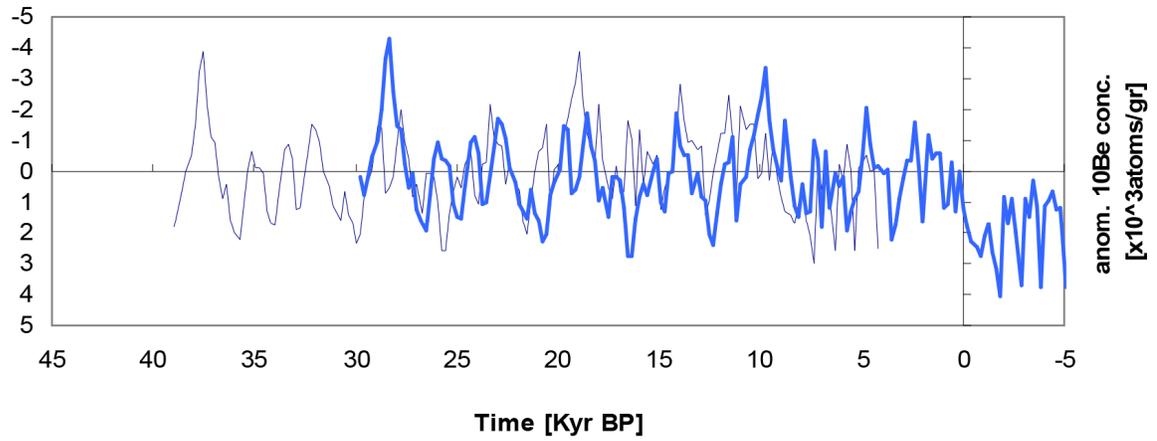
2 Figure S2. GISP2 icecore analysis. a) 10Be concentration [N97], b) ice accumulation rate [Alley, 2000], c) linear
 3 correlation between 10Be concentration and ice accumulation shown in a) and b).



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2 Figure S3. GISP2 icecore analysis. a) 10Be residual concentration after to eliminate ice accumulation linear
 3 contributions (Fig S4c), b) Standard deviation of a), c) Demodulated 10Be [(a)/(b)] with a precession simple
 4 model with linear and sinusoidal variation for detrending.

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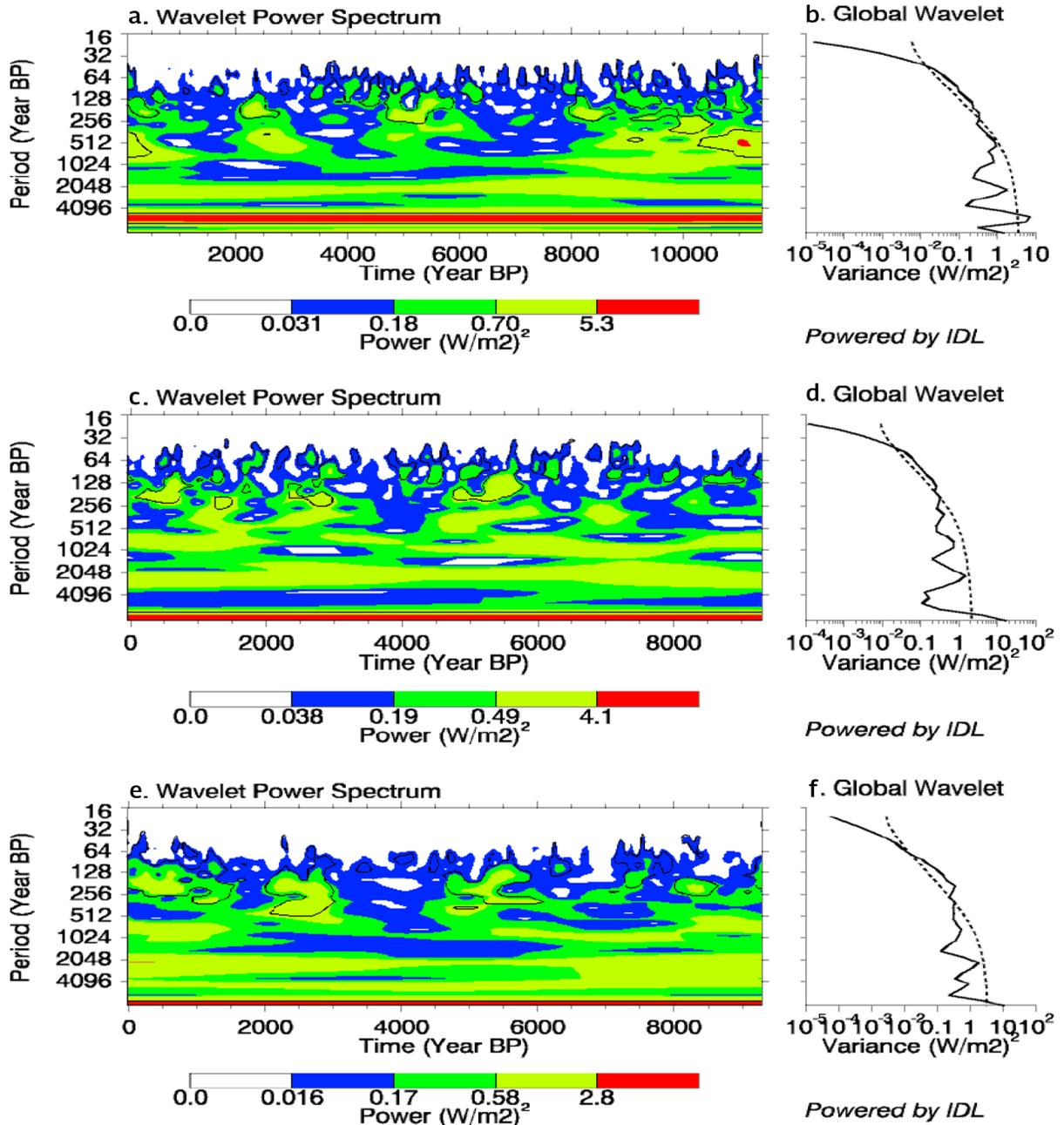
2 Figure S4. GISP2 icecore analysis. 10Be detrended and demodulated residual concentration (DDR) after to: a)
3 subtract ice accumulation linear contributions (Fig S4c), b) divide by standard deviation (Fig. S5b), c) subtract
4 linear and sinusoidal variation (Fig S5c). A lagged model (9.4 Kyr) of this DDR of 10Be is also displayed.

5

1 **SI-4. Spectral analysis of solar activity**

2 In order to verify the multi-millennia scale solar oscillation of ~9500 yrs detected in 10Be
3 records, a wavelet analysis was applied to the reconstructed records of TSI, using the online
4 resource by Torrence and Compo (1998).

5 TSI spectral results, displayed in Figure S1, show three main, significant (~10% level)
6 periodicities around 9000, 5000, and 2400 years. It should be noted that the *S09* record also
7 shows relatively high-frequency significant variability with periods around 200 years.



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2 Fig. S5: (a, c, and e.) The wavelet power spectrum of TSI(S04), TSI(S09), and TSI(S12), respectively. The
 3 contour levels are chosen so that 75%, 50%, 25%, and 5% of the wavelet power is above each level,
 4 respectively. Black contour is the 10% significance level, using a red-noise (autoregressive lag1)
 5 spectrum. (b, d, and f.) The global wavelet power spectrum (black line). The dashed line is the significance for
 6 the global wavelet spectrum, assuming the same significance level and background spectrum as in (a, c, and e.,
 7 respectively). Reference: Torrence, C. and G. P. Compo, 1998: A Practical Guide to Wavelet Analysis. Bull.
 8 Amer. Meteor. Soc., 79, 61-78.